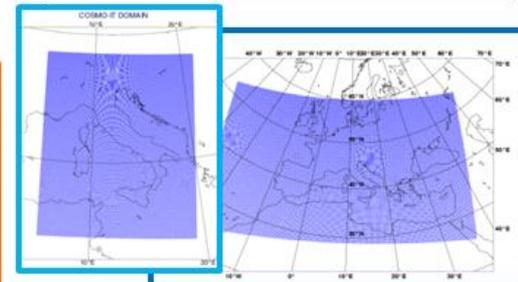
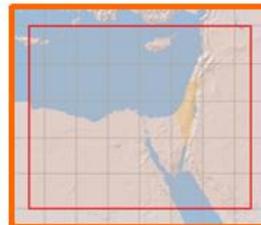
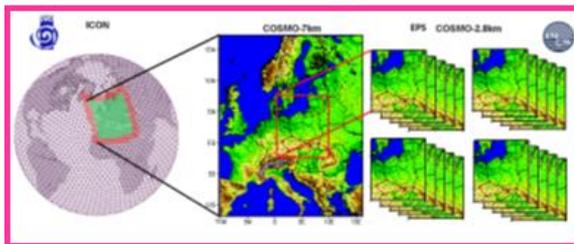
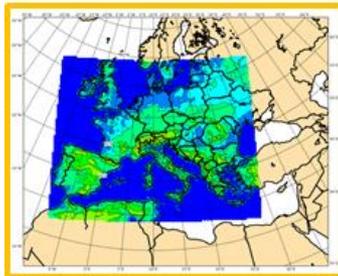
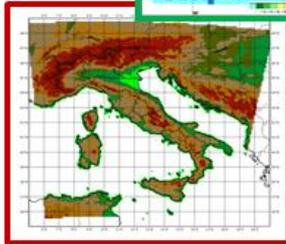


# Model perturbations for the COSMO ensembles

Chiara Marsigli  
Deutscher Wetterdienst

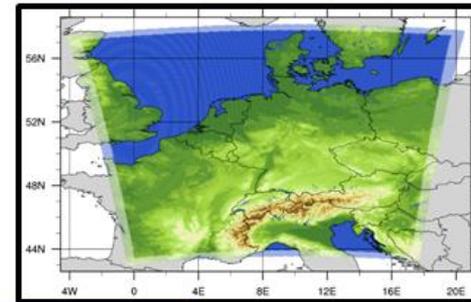
## Outline

- The COSMO ensembles
- Recent developments:
  - Sensitivity of ICON to physics parameters
  - Combination of SPPT and PP and in COSMO
  - More stochastic physics!
- Use of ensembles
- Final remarks



## Ensemble systems

- ICON-D2-EPS
- COSMO-2E  
COSMO-1E
- TLE-MVE
- COSMO-2I-EPS
- COSMO-IT-EPS
- COSMO-Ru2-EPS
- COSMO-IL-ENS
- COSMO-LEPS
- COSMO-ME-EPS



<http://www.cosmo-model.org/content/tasks/workGroups/wg7>

## Model perturbation

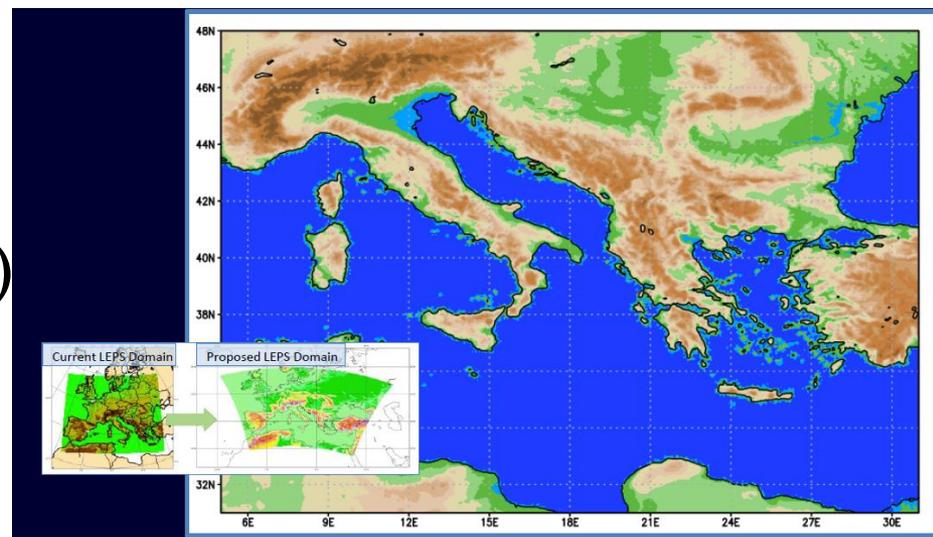
- Methods currently operational:
  - SPPT - Stochastically Perturbed Parametrization Tendency
  - PP – Perturbed Parameters
- Methods currently under development:
  - SMME – Stochastic Model for the Model Error
  - AMPT – Additive Model-error perturbations scaled by Physical Tendencies
- New (external) developments currently under test:
  - SSC - Stochastic shallow convection
  - PSP2 - Physically based stochastic perturbations for boundary layer turbulence

## Status of the experiments on parameter perturbations towards ICON-LEPS

Euripides N. Avgoustoglou  
Hellenic National Meteorological Service

- Goal: study the sensitivity of the ICON model to a large set of parameters over a Mediterranean area

=> implementation of the  
Parameter Perturbation  
in ICON-LEPS  
(transition of COSMO-LEPS)



E. Avgoustoglou, HNMS

## Status of the experiments on parameter perturbations towards ICON-LEPS

Euripides N. Avgoustoglou  
Hellenic National Meteorological Service

24 parameters were considered.



3 values/parameter including default.



The evaluation period consisted of 62 days from year 2020 i.e.:  
January 1-31, July 1-31

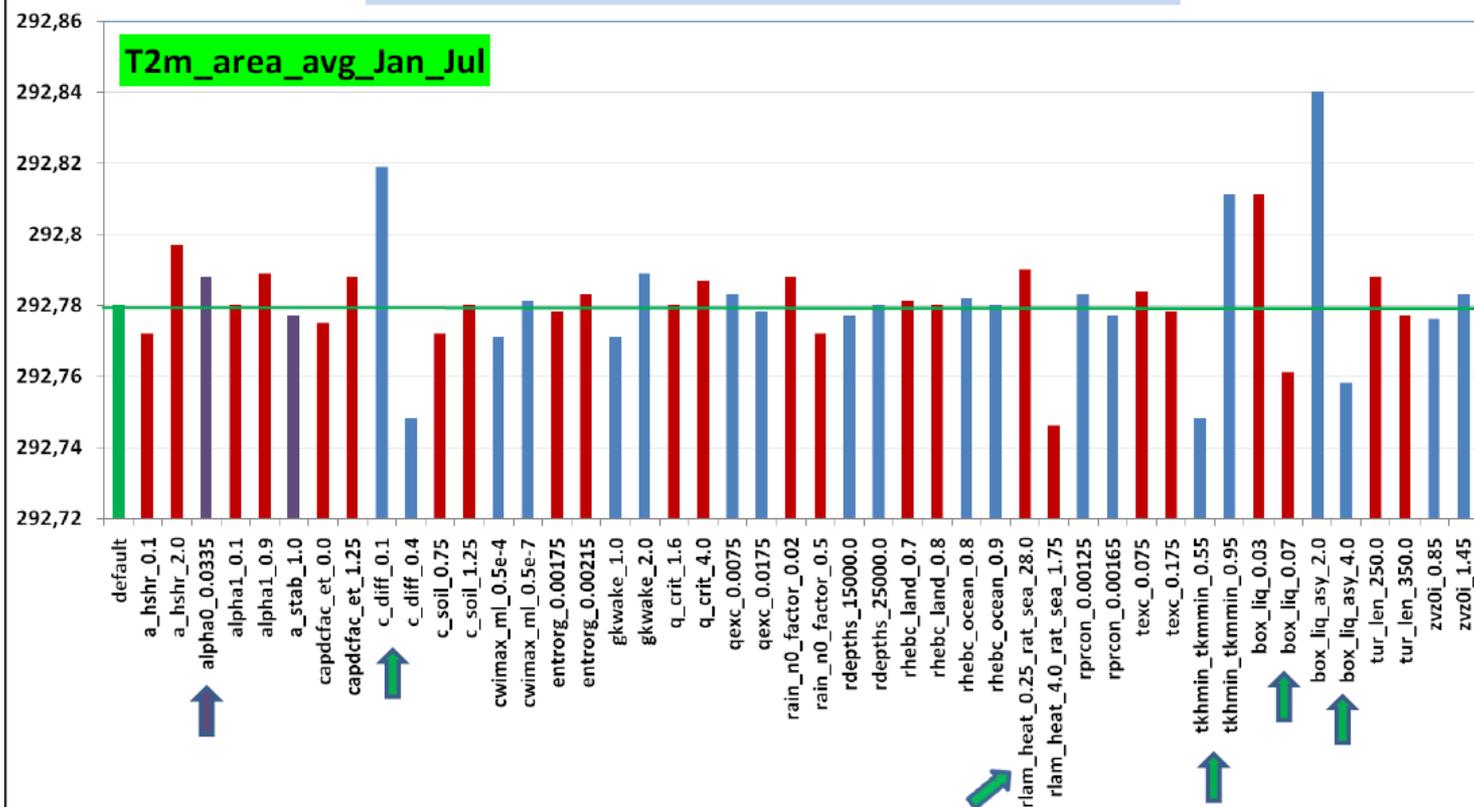
3000 runs based on ICON-IMS (Gratis IMS):

- ⊕ Horizontal grid size: R3B08 (~6.5km).
- ⊕ 417x273 grid points (wider area of Greece and Italy), 65 levels.
- ⊕ Integration time-step: 60 secs.
- ⊕ Integration period: 132 hs.
- ⊕ Boundary conditions : 3hr IFS Forecast.
- ⊕ Computational Cost ~  $5 \times 10^6$  b.u. on Cray X C40 of ECMWF (Gratis HNMS).

E. Avgoustoglou, HNMS



### Areal <2m temperature> [K] (132<sup>nd</sup> hr)



Euripides Avgoustoglou, Hellenic National Meteorological Service, 23<sup>rd</sup> COSMO GM, September 13<sup>th</sup> 2021

07

### Dominating parameters table for the considered meteorological fields



T2m	2 m Temperature [K]	box_liq_asy	c_diff	tkhmin_tkmmin
Tmax2m	2 m max Temperature [K]	box_liq_asy	c_diff	tkhmin_tkmmin
Tmin2m	2 m min Temperature [K]	tkhmin_tkmmin	rlam_heat_0.25_rat_sea_28.0	c_diff
Td2m	2 m max dew point Temperature [K]	c_diff	box_liq	alpha1
tot_prec	accumulatedPrecipitation [kg/m^2]	box_liq_asy	rain_n0_factor	rprcon
pmsl	mean sea level Pressure [Pa]	box_liq	gkwake	c_diff
u10m	10 m wind speed u component [m/s]	box_liq	box_liq_asy	tur_len
v10m	10 m wind speed v component [m/s]	gkwake	tkhmin_tkmmin	box_liq_asy
gust10m	wind gust 10 m above ground [m/s]	gkwake	box_liq	a_hshr
clcl	low cloud cover [1-100]	box_liq_asy	box_liq	alpha0
clcm	medium cloud cover [1-100]	box_liq	box_liq_asy	rdepths
clch	high cloud cover [1-100]	zvz0i	entrorg	rprcon
clct	total cloud cover [1-100]	box_liq_asy	box_liq	zvz0i
tqv	column integrated water vapour [kg/m2]	alpha1	a_stab	box_liq
tqi	total column integrated cloud ice [kg/m2]	zvz0i	rprcon	box_liq_asy
tqc	total column integrated cloud water [kg/m2]	rdepths	tur_len	q_crit

Euripides Avgoustoglou, Hellenic National Meteorological Service, 23<sup>rd</sup> COSMO GM, September 13<sup>th</sup> 2021

23

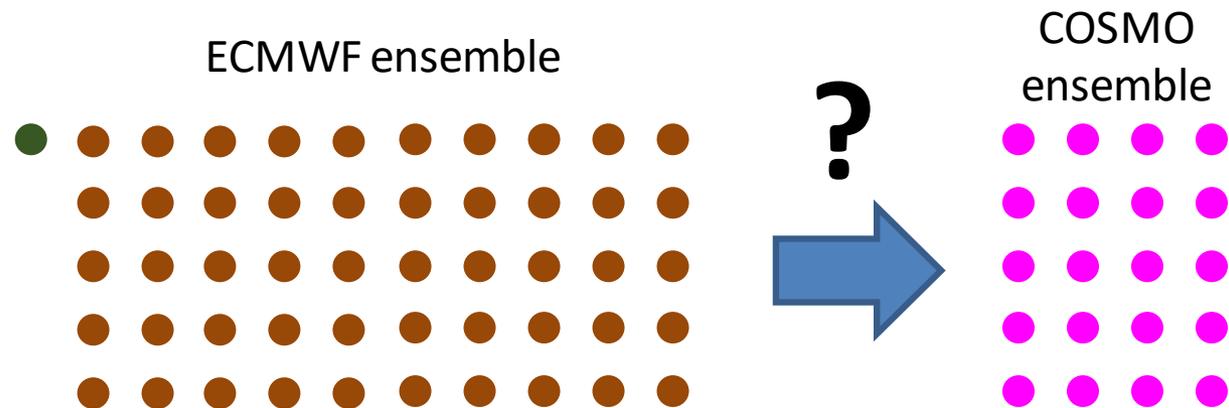
## COSMO-IL-ENS

Sources of forecast uncertainty:

- Uncertainty in boundary conditions → use of driving ensemble (**EC-ENS**)
- Uncertainty in model physics → Stoch. Pert. of Param. Tendencies (**SPPT**), parameter perturbations (**PP**)

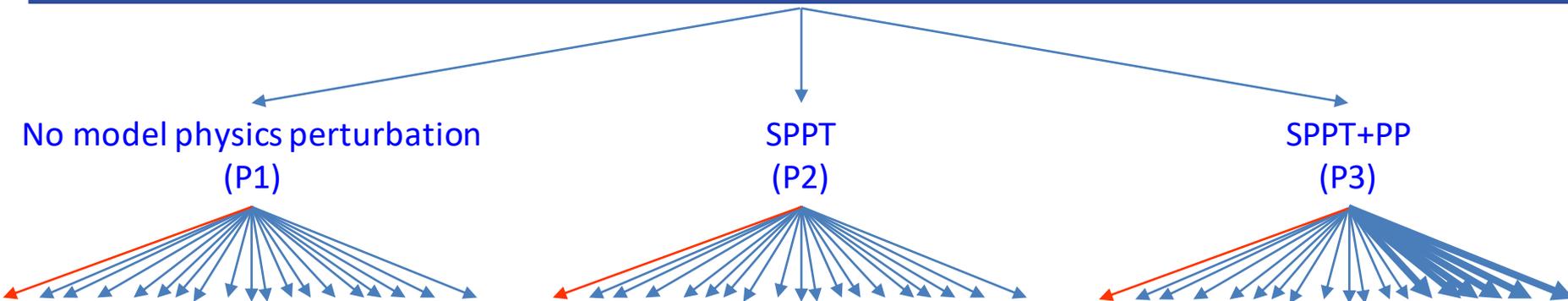
### Questions?

1. Which 20 EC-ENS members to choose?
2. Do model physics perturbations benefit?

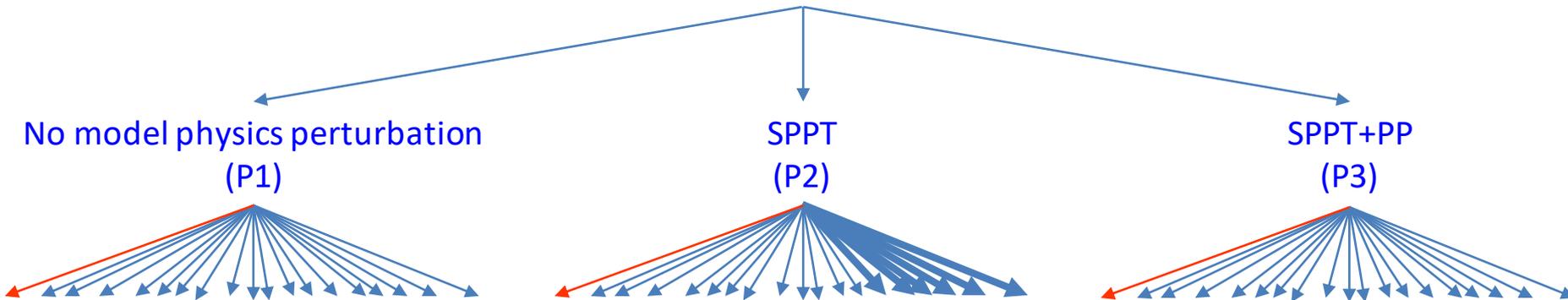


Khain et al., IMS

# Which option of ensemble structure is better for **near surface variables**?

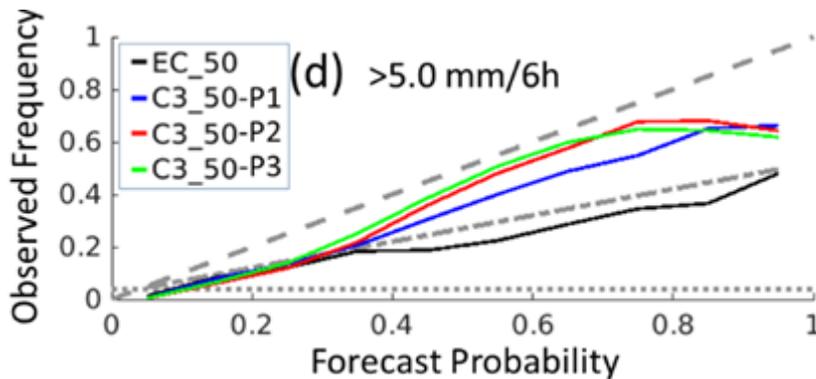


# Which option of ensemble structure is better for **precipitation**?



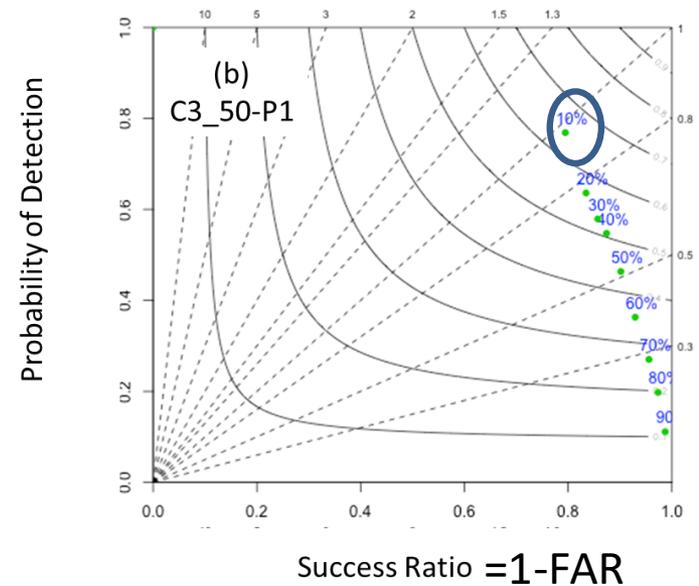
Reliability: P3 (SPPT+PP), Resolution: P1 (no phys pert), ROC and TS: P2 (SPPT)

Reliability diagram



C3\_50-P3 has better reliability

Performance diagram



Problem: The ensemble is biased. One should forecast the event (>5mm/6h) already when 10% of members “agree” with that.

C3\_50-P1 has a bit higher TSmax



C3\_50-P2

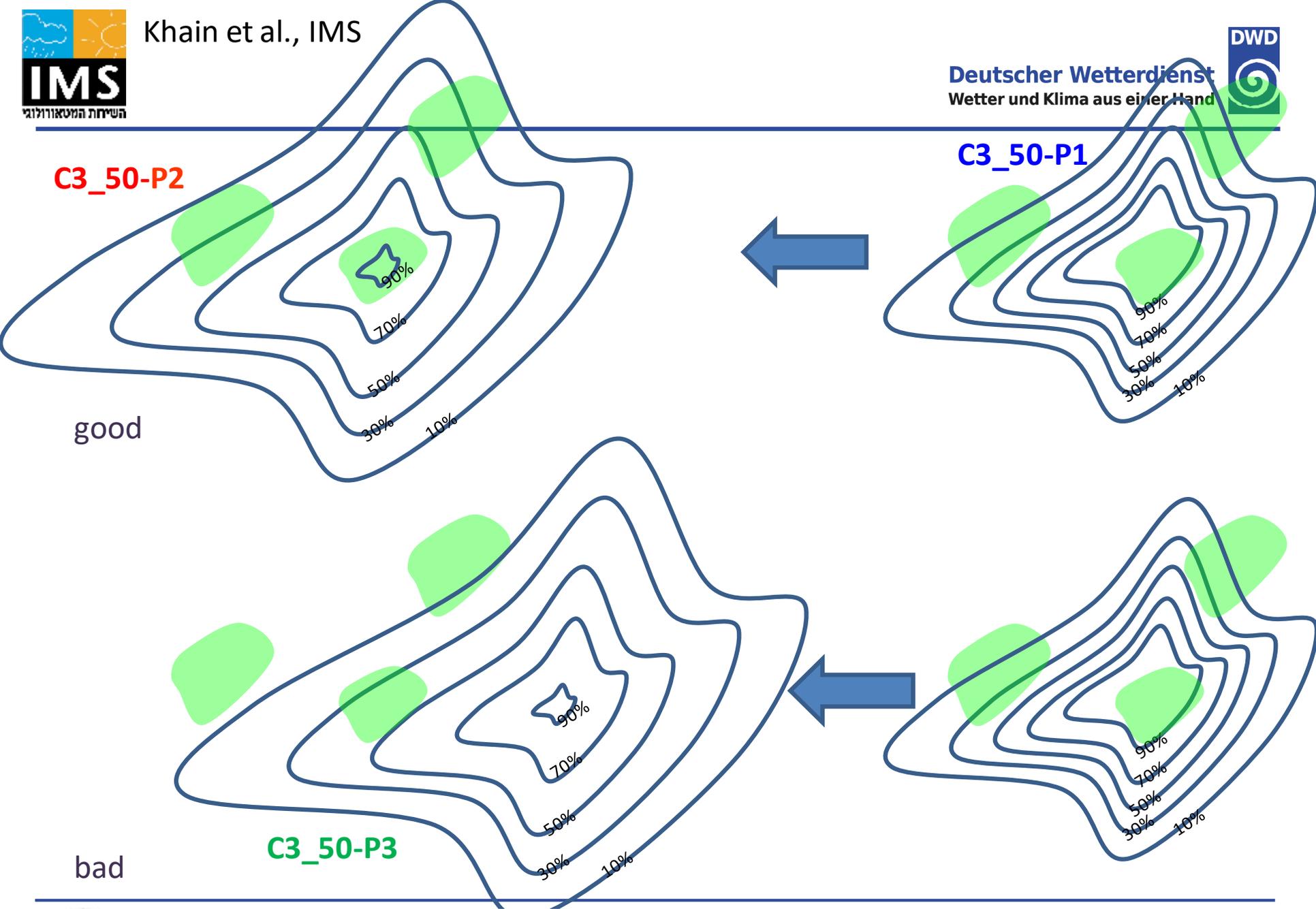
C3\_50-P1

good

bad

C3\_50-P3

Increasing spread is good if it does not destruct the model (precipitation locations etc.)



## Stochastic Workshop

- 2nd and 3rd of March 2021
- Purpose: to make the point about the activities on-going in the COSMO Consortium and in the other European Consortia in the field of "stochastic physics", in particular intrinsically stochastic parametrisations, in view of their usage in ensembles
- 40-50 participants, from COSMO members, LMU (University of Munich), ECMWF, Meteo France, Met Office, Met Eireann, KNMI, SMHI, Met No, Met Hu, AEMET, NCAR, KIT
- Presentations and minutes (of a very interesting discussion) are online at: <http://www.cosmo-model.org/content/tasks/workGroups/wg7/default.htm>

Physically based stochastic perturbations for boundary layer turbulence :  
PSP (Kober and Craig, 2016)                      PSP2 (Hirt et al., 2019, MWR)

It reintroduces the influence of the lost small-scale variability by adding perturbations to the tendencies of  $T, q_v, w$  on the smallest effectively resolved scale ( $5\Delta x$ )

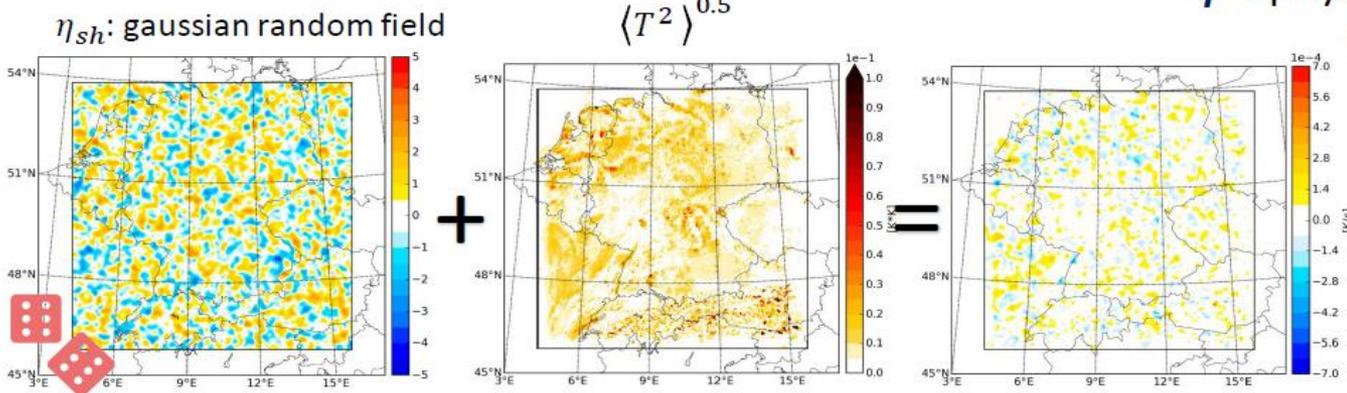
$$\left(\frac{\partial \phi}{\partial t}\right)_{all} = \frac{\partial \phi}{\partial t} + \underbrace{\alpha \cdot \eta \cdot \sqrt{\overline{\phi'^2}}}_{\text{Stochastic perturbations}}$$

$$\phi = \{T, q, w\}$$

$\eta(t, \sigma)$ : Random field , regenerated every 10 min with spatial correlation  $\sigma$

$\alpha$ : perturbation ampl., scaling factors

$\phi'$ : physical scaling/subgrid-scale variance of variable  $\phi$



(Kober and Craig, 2016)

## Physically based stochastic perturbations for boundary layer turbulence (PSP2)

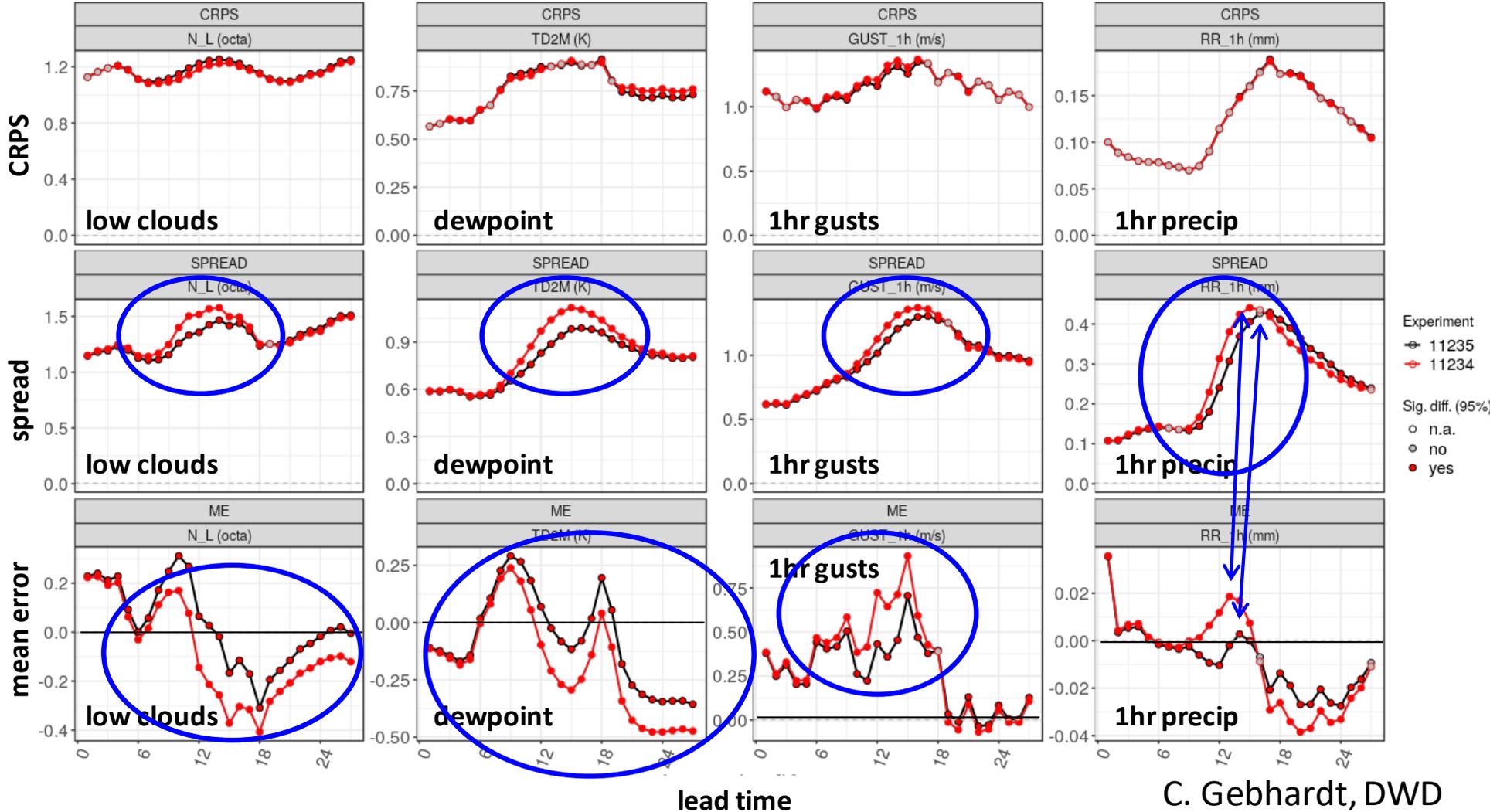
- cooperation with Ludwigs-Maximilian-Universität in Munich (LMU)
- first promising tests at LMU
- PSP2 implemented in ICON by LMU
- **test run of ICON-D2-EPS from May 26<sup>th</sup> to August 1<sup>st</sup> 2021 at DWD**
- ICON-D2-EPS with PSP2 vs. operational set up
- verification against synop observations



C. Gebhardt, C. Marsigli, DWD

2021/05/26 22UTC - 2021/08/01 21UTC  
 INI: 00 UTC, DOM: ALL

**PSP2**      **oper.**      **00 UTC runs**



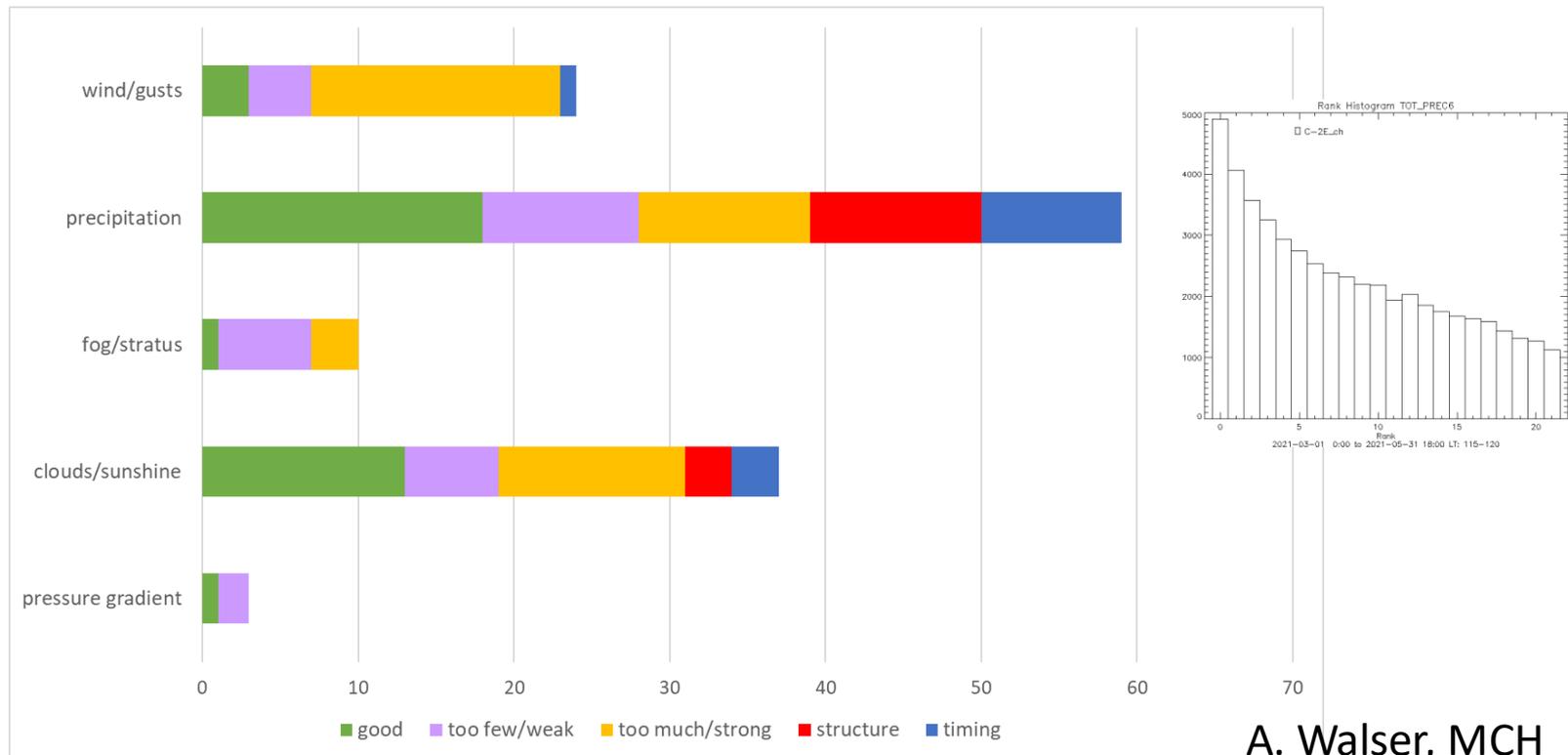
C. Gebhardt, DWD



# Forecaster feedbacks 2021



- model feedbacks from forecaster on duty every day (scheduled, up to 15 min)
- increased use of ensemble information also in the short range
- probabilities are translated to keywords in forecast bulletins (possible, likely, ...)
- ...but control run still get (too) much attention

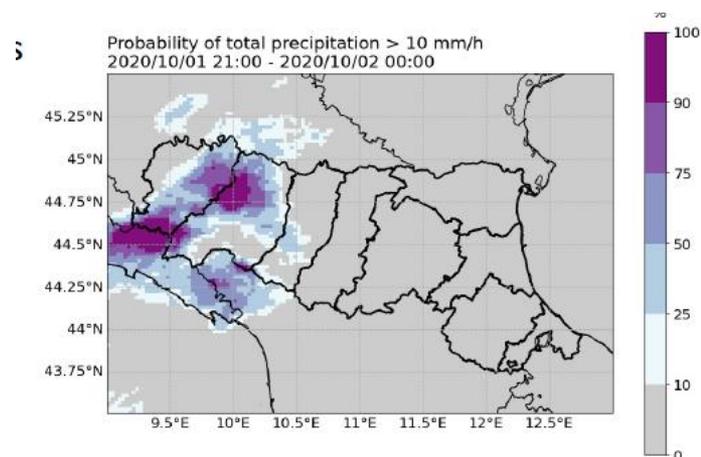
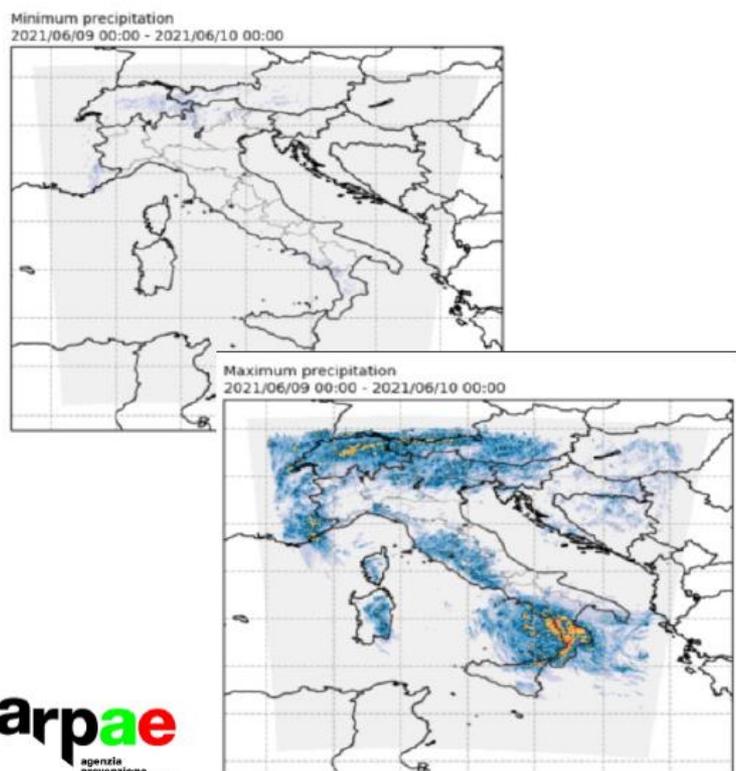


A. Walser, MCH



# COSMO-2I-EPS

- Most likely scenario (mean)
- Extreme scenarios (min max)
- Probability maps
- Percentiles

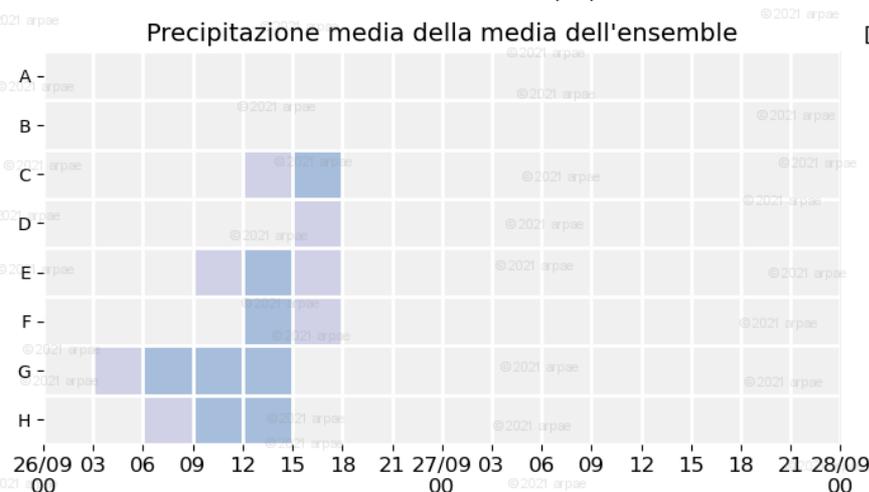
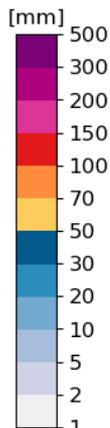


C. Marsigli, V. Poli et al., Arpae-SIMC

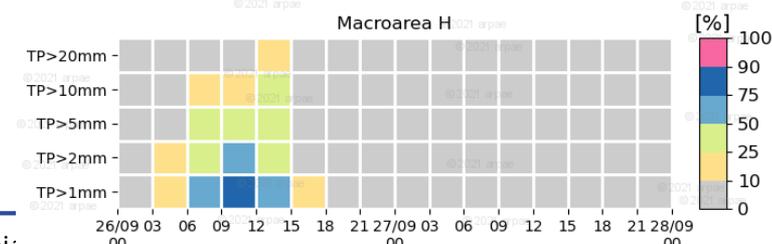
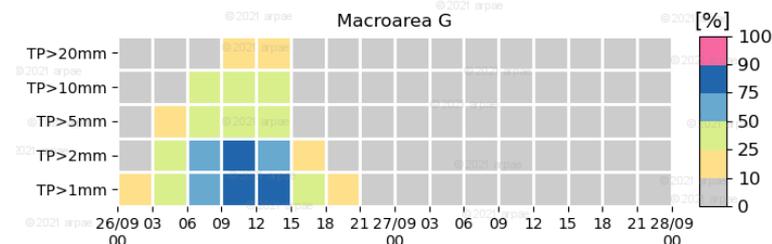
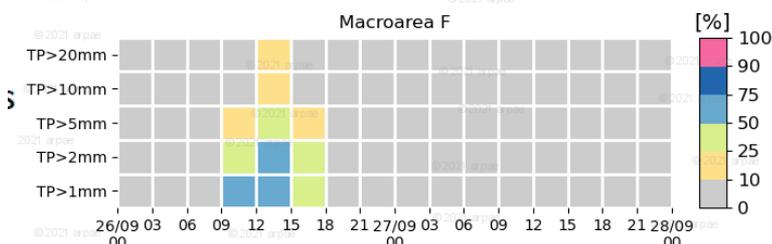
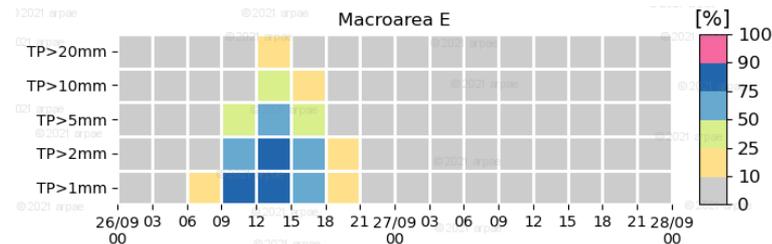
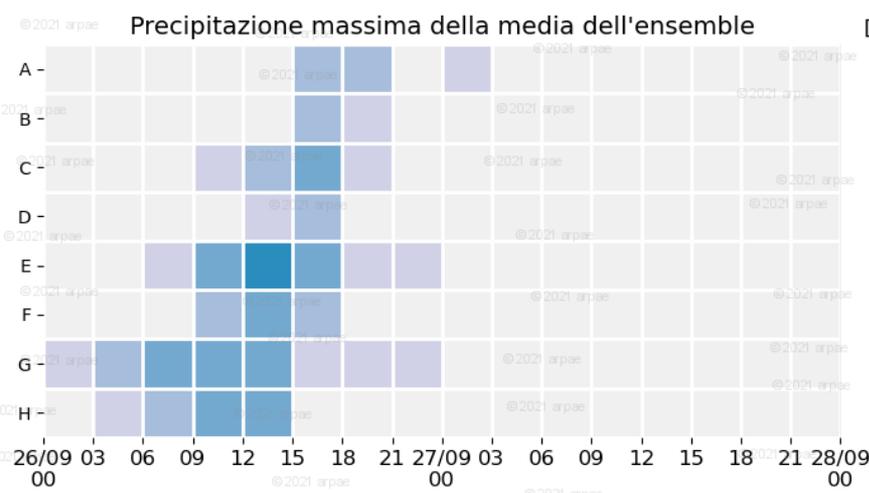
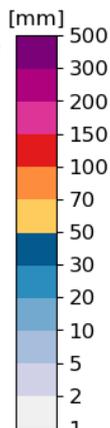
# COSMO-2I-EPS

Corsa di COSMO-2I-EPS del 25/09/2021 21:00

Precipitazione media della media dell'ensemble



Precipitazione massima della media dell'ensemble



Chi:



## Final remarks

- The transition of the ensembles to ICON is on-going
- Model perturbation schemes are further studied / developed / tested
- The ensemble development is becoming more and more part of the numerical modelling development
- It is needed to invest more in ensemble interpretation, in order to increase its usability and usage

**Thank you for your attention!**