



Multilevel Time Integrators for Large-scale Atmospheric Flows

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SRNWP Workshop, Deutscher Wetterdienst, Offenbach, Germany
May 13, 2013

Acknowledgements



Dahlem Research School



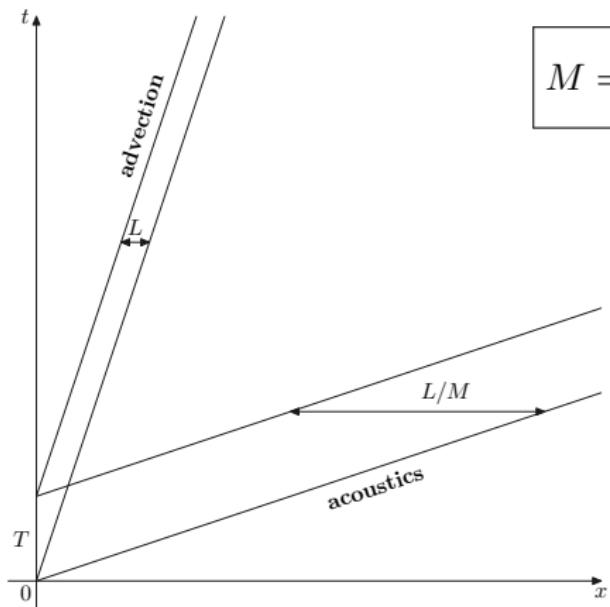
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Atmosphere dynamics: different **spatial** and **temporal** scales



Atmosphere dynamics: different **spatial** and **temporal** scales



$$M = \frac{U}{c}$$

For atmospheric flows

$$0 < M \lesssim 0.3$$



Sound waves: physically **unimportant**, computationally **demanding**

For **explicit** methods, **CFL**

$$\frac{c\Delta t}{\Delta x} < 1$$

Analytical approach: **reduced** models

Anelastic-like models on small scales, **hydrostatic** on large scales



Numerical approach: semi-implicit fully compressible models

The **implicit** part handles **sound** and/or **gravity modes**

No stability constraints, but dispersive



A conservative, semi-implicit, fully compressible model



A conservative, semi-implicit, fully compressible model

- ▶ **Second-order** accurate throughout
- ▶ **Consistent** with soundproof models
- ▶ With **selective** control over resolved scales/effects



Compressible Euler Equations

$$\rho_t + \nabla \cdot (\rho \mathbf{v}) = 0$$

$$(\rho \mathbf{v})_t + \nabla \cdot (\rho \mathbf{v} \otimes \mathbf{v}) + \nabla p = -\rho g \mathbf{k}$$

$$\alpha P_t + \nabla \cdot (P \mathbf{v}) = 0$$

Equation of state, ideal gas:

$$P = \rho \theta = \frac{p_{ref}}{R} \left(\frac{p}{p_{ref}} \right)^{1/\gamma}$$

Initial and boundary conditions

Durran's pseudo-incompressible model: $\alpha = 0, P = \bar{P}(z)$



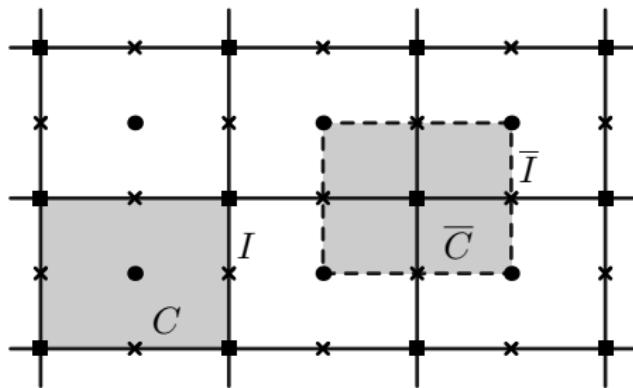
Fractional step method (Klein (2009)):

- ▶ **Predictor:** auxiliary hyperbolic system
- ▶ **Corrector:** two projections for non-compliant fluxes and use of p^n



Fractional step method (Klein (2009)):

- ▶ Predictor: auxiliary hyperbolic system
- ▶ Corrector: two projections for non-compliant fluxes and use of p^n





Finite volume scheme in **conservation** form

Time integration: second-order SSP Runge-Kutta, $\text{CFL}_{\text{adv}} < 1$

$$(\rho \mathbf{v})_C^{n+1,*} = (\rho \mathbf{v})_C^n - \frac{\Delta t}{|C|} \left(\int_{\partial C} \rho \mathbf{v} (\mathbf{v} \cdot \mathbf{n}) \, dl \right)^{n+\frac{1}{2}} - \frac{\Delta t}{|C|} \left(\int_{\partial C} p \cdot \mathbf{n} \, dl \right)^n$$



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- ▶ Velocities **not** compliant with:

$$[\alpha P_t + \nabla \cdot (P \mathbf{v})]_{n+\frac{1}{2}} = 0$$

- ▶ Old pressure used in the momentum equation



Corrector step – First projection

Integrate $\rho \mathbf{v}$ over a **half** time step, replace in

$$[\alpha P_t + \nabla \cdot (P\mathbf{v})]_{n+\frac{1}{2}} = 0$$

with

$$\frac{\partial P}{\partial t} \approx \frac{\partial P}{\partial p} \frac{\delta p}{\Delta t}$$

Corrector step – First projection

Integrate ρv over a **half time step**, replace in

$$[\alpha P_t + \nabla \cdot (P\mathbf{v})]_{n+\frac{1}{2}} = 0$$

with

$$\frac{\partial P}{\partial t} \approx \frac{\partial P}{\partial p} \frac{\delta p}{\Delta t}$$

Helmholtz equation

$$\alpha C_1 \delta p + \nabla \cdot (C_2 \nabla \delta p) = \nabla \cdot \left[(P\mathbf{v})^{n+\frac{1}{2}} \right]$$

Equivalent to **wave** equation for p , consistent with **soundproof** limit

$$C_1 \propto \frac{M^2}{CFL_{adv}^2}$$



Corrector step – Second projection

On the **dual grid**, integrate ρv over a **full time step**, replace in

$$\frac{P^{n+1} - P^n}{\Delta t} + \frac{1}{2} \nabla \cdot \left[(P\mathbf{v})^{(n+1)} + (P\mathbf{v})^{(n)} \right] = 0$$

TRAP**or**

$$\frac{\frac{3}{2}P^{n+1} - 2P^n + \frac{1}{2}P^{n-1}}{\Delta t} + \nabla \cdot (P\mathbf{v})^{(n+1)} = 0$$

BDF2



Corrector step – Second projection

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TRAP

or

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BDF2

Second **Helmholtz** equation

$$\alpha C_1 \delta p_n + \nabla \cdot (C_2 \nabla \delta p_n) = RHS$$

Momentum and pressure corrected, full second-order scheme



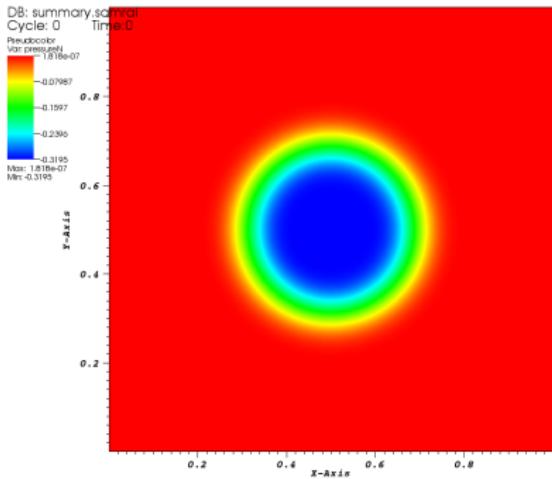
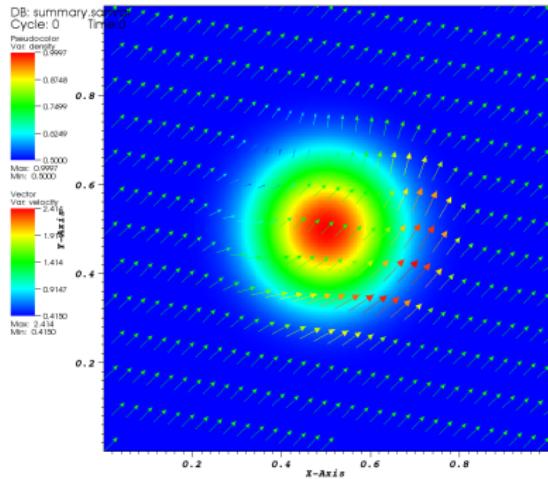
C++ code, SAMRAI & Hypre

Overhead of compressibility: added diagonal term

Results

Traveling rotating vortex in $[0, 1]^2$ m, Kadioglu et al. (2008)

Exact solution in the incompressible case, $(u_{bg}, v_{bg}) = (1, 1)$



user: benacchio
Wed Oct 3 18:08:46 2012

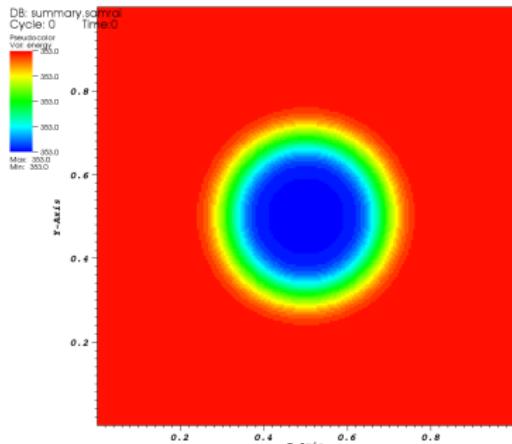
user: benacchio
Tue Oct 2 20:23:15 2012

$$\rho_{ref} = 0.5 \text{ kg/m}^3, p_{ref} = 101625 \text{ Pa}, T_{ref} = 706.098 \text{ K}$$

$$M \approx 4.5 \text{e-}03$$

Results

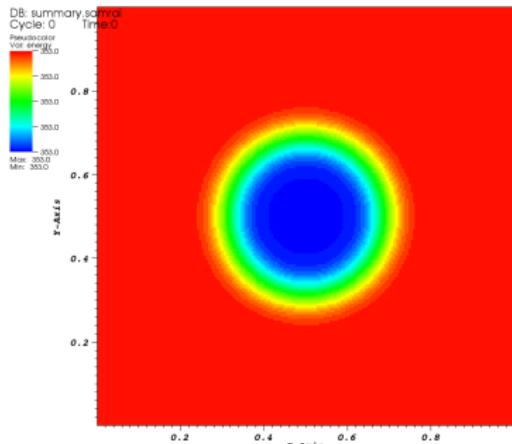
$P_{t=0}$ derived from $p_{t=0}$ through EOS



user: benacchio
Tue Oct 2 19:51:58 2012

Results

$P_{t=0}$ derived from $p_{t=0}$ through EOS



192×192 cells

$$\Delta t = CFL_{adv} \Delta x / U_{max}$$

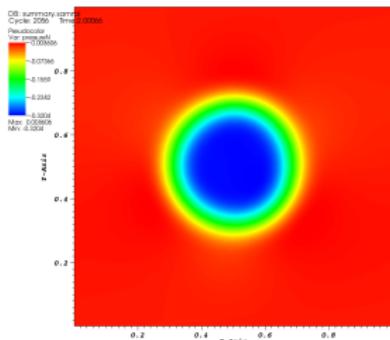
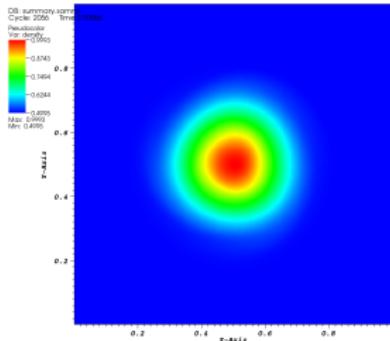
$$CFL_{adv} = \boxed{0.45}$$

$$CFL_{ac} \approx \boxed{99.3}$$

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Tue Oct 2 19:51:58 2012

$T = 2\text{ s}$

COMPRESSIBLE

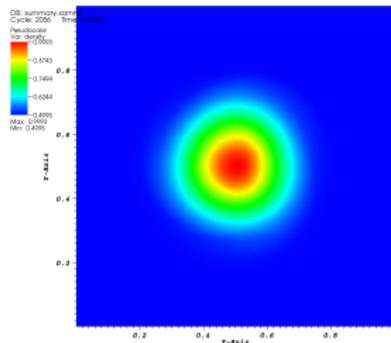


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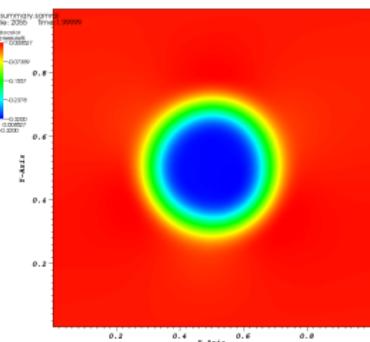
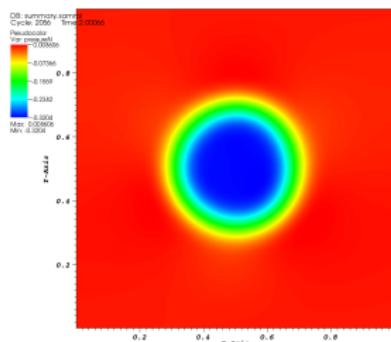
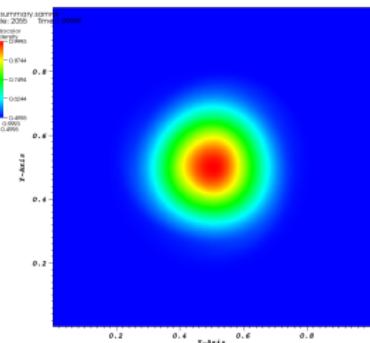
$T = 2 s$



COMPRESSIBLE

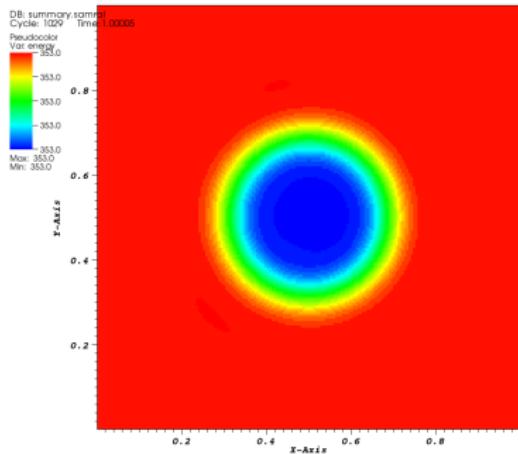


PSEUDO-INCOMPRESSIBLE

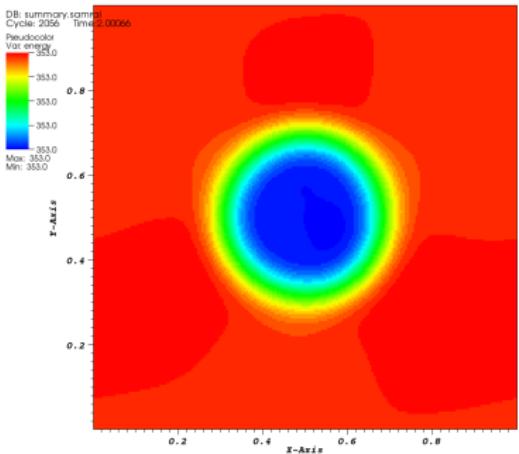


Transport of P

$T = 1\text{ s}$



$T = 2\text{ s}$



user: benacchio
Tue Apr 23 12:29:54 2013

user: benacchio
Tue Apr 23 12:30:07 2013

$$\alpha P_t + \nabla \cdot (P\mathbf{v}) = 0$$

Results

Simple acoustic wave

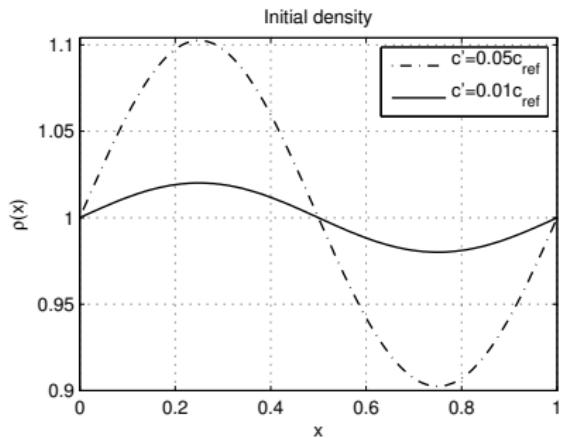
$$u = \frac{2}{\gamma - 1} c' \sin \left(\frac{2\pi}{L} x \right)$$

$$\rho = \rho_{ref} \left[1 + \frac{c'}{c_{ref}} \sin \left(\frac{2\pi}{L} x \right) \right]^{\frac{2}{\gamma-1}}$$

$$p = p_{ref} \left[1 + \frac{c'}{c_{ref}} \sin \left(\frac{2\pi}{L} x \right) \right]^{\frac{2\gamma}{\gamma-1}}$$

Rieper(2011), Vater(2013)

$x \in [0, 1]$, periodic BCs

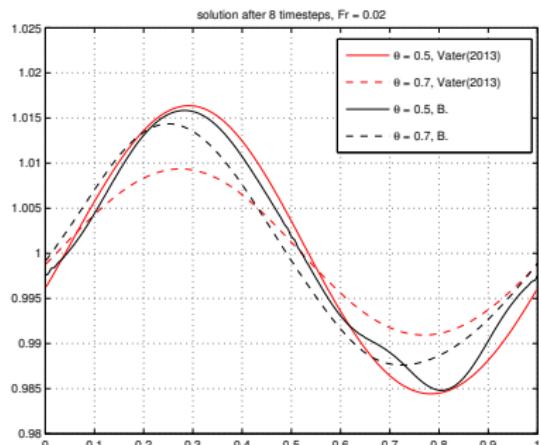
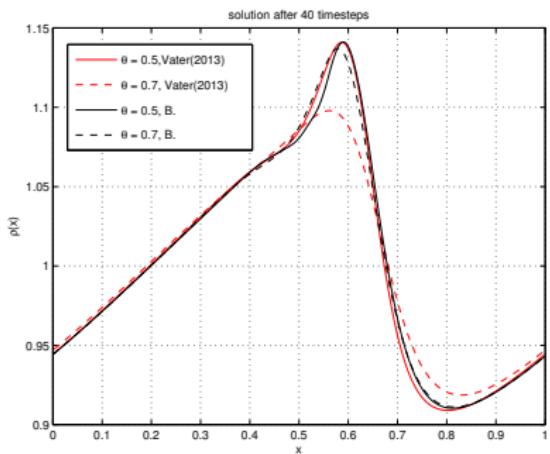


$$\rho_{ref} = 1 \text{ kg/m}^3, \gamma = 2$$

$$T_{ref} \approx 353 \text{ K}, c_{ref} \approx 450 \text{ m/s}$$

Results

$$\Delta x = 1/256 \text{ m}, \text{ CFL}_{\text{adv}} = 0.77$$



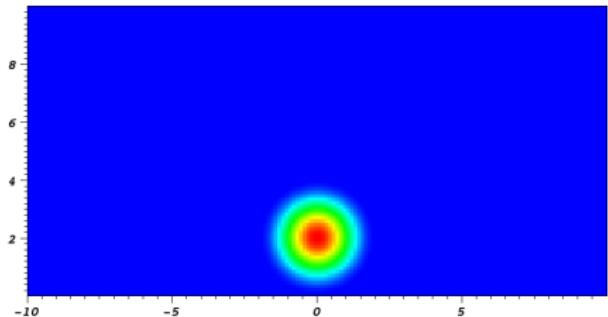
$$c' = 0.05 c_{\text{ref}}, M(t=0) \approx 0.1$$

$$c' = 0.01 c_{\text{ref}}, M(t=0) \approx 0.02$$

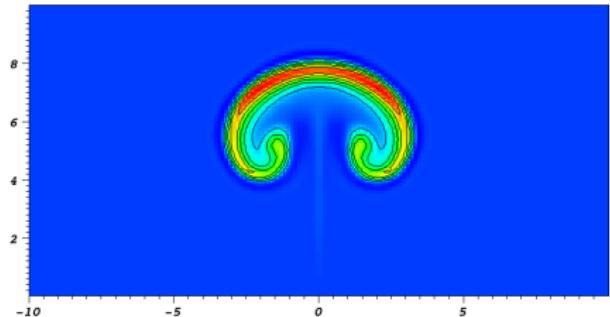
Results

**Warm air bubble in
[−10, 10] × [0, 10] km neutrally
stratified atmosphere, Klein(2009)**
BDF2, CFL_{adv} = 0.5, Δx = 125 m

$$\theta'(x, z) = \begin{cases} 2K \cos^2\left(\frac{\pi}{2}r\right) & (r \leq 1) \\ 0 & \text{otherwise} \end{cases}$$



T = 0 s



T = 1000 s, $\delta\theta'_{\max} \approx 1.7 \text{ K}$

Results

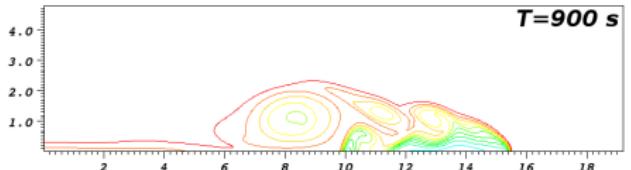
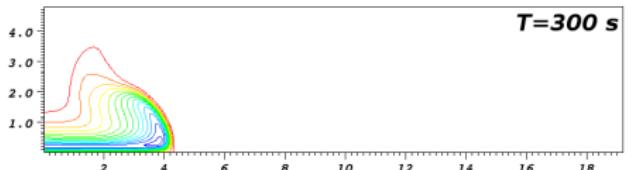
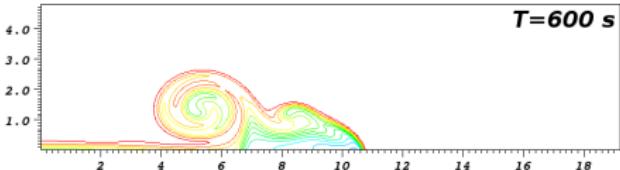
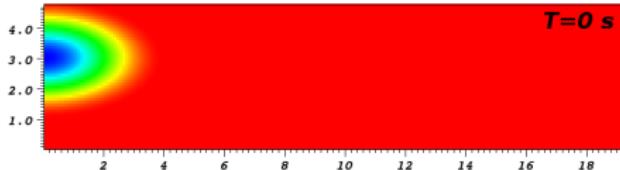
Density current

Straka et al. '93, Mueller et al. '12

BDF2, CFL_{adv} = 0.65, $\Delta x \approx 28 \text{ m}$

$$T' = \begin{cases} -15 \text{ K} \left[\frac{1}{2}(1 + \cos(\frac{\pi}{2}r)) \right] & (r \leq 1) \\ 0 & \text{otherwise} \end{cases}$$

$$+ \mu = 75.0 \text{ m}^2/\text{s}$$



Front position: $x = 15497.71 \text{ m}$ (M. '12: 15452.4 m)

$|\theta'_{\max}| = 10.13 \text{ K}$ (M. '12: 9.8 K)



Fully compressible, semi-implicit model for the simulation of multiscale atmospheric processes



Fully compressible, semi-implicit model for the simulation of multiscale atmospheric processes

- ▶ Compatible with pseudo-incompressible soundproof model
- ▶ Conservative and second-order accurate
- ▶ Platform to investigate processes interactions at large scales



Multilevel time discretization

Vater (2013): **multigrid** in time, blend of **implicit** schemes

Damping and **dispersion** properties used to eliminate **unresolved modes** and transport long waves **accurately**

Multiscale features of **large-scale** test cases modelled with **controlled treatment** of acoustic/gravity waves





Corrector step – First projection

Integrating $\rho \mathbf{v}$ over a **half time step**

$$\begin{aligned}(\rho \mathbf{v})^{n+\frac{1}{2}} &= (\rho \mathbf{v})^n - \frac{\Delta t}{2} [\nabla \cdot (\rho \mathbf{v} \otimes \mathbf{v})]^{n+\frac{1}{4}} - \frac{\Delta t}{2} \nabla (p^n + \delta p) \\&= (\rho \mathbf{v})^{n+\frac{1}{2},*} - \frac{\Delta t}{2} \nabla \delta p\end{aligned}$$

Corrector step – First projection

Integrating $\rho\mathbf{v}$ over a half time step

$$\begin{aligned}(\rho\mathbf{v})^{n+\frac{1}{2}} &= (\rho\mathbf{v})^n - \frac{\Delta t}{2} [\nabla \cdot (\rho\mathbf{v} \otimes \mathbf{v})]^{n+\frac{1}{4}} - \frac{\Delta t}{2} \nabla (p^n + \delta p) \\&= (\rho\mathbf{v})^{n+\frac{1}{2},*} - \frac{\Delta t}{2} \nabla \delta p\end{aligned}$$

With

$$\frac{\partial P}{\partial t} = \frac{\partial P}{\partial p} \frac{\partial p}{\partial t} \approx \frac{R^{-\gamma}}{\gamma} \left(\frac{P}{p_{ref}} \right)^{1-\gamma} \frac{\delta p}{\Delta t}$$

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the constraint

$$[\alpha P_t + \nabla \cdot (P\mathbf{v})]_{n+\frac{1}{2}} = 0$$



Corrector step – First projection

Integrating $\rho\mathbf{v}$ over a half time step

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With

$$\frac{\partial P}{\partial t} = \frac{\partial P}{\partial p} \frac{\partial p}{\partial t} \approx \frac{R^{-\gamma}}{\gamma} \left(\frac{P}{p_{ref}} \right)^{1-\gamma} \frac{\delta p}{\Delta t}$$

the constraint

$$[\alpha P_t + \nabla \cdot (P\mathbf{v})]_{n+\frac{1}{2}} = 0$$

yields a Helmholtz equation:

$$-\frac{\alpha}{\gamma \Delta t} \left[P^{n+\frac{1}{2},*} \right]^{1-\gamma} R^{-\gamma} p_{ref}^{\gamma-1} \delta p + \nabla \cdot \left[\frac{\Delta t}{2} \left(\frac{P}{\rho} \right)^{n+\frac{1}{2},*} \nabla \delta p \right] = \nabla \cdot \left[(P\mathbf{v})^{n+\frac{1}{2},*} \right]$$

Corrector step – First projection

If $\delta p = p^{n+1} - p^n$, equivalent formulation with $\alpha = 1$

$$-\frac{2}{\Delta t^2 c^2} (p^{n+1} - p^n) + \nabla^2 p^{n+1} = \frac{2}{\Delta t} \rho^{n+\frac{1}{2},*} \nabla \cdot \mathbf{v}^n$$

- ▶ If $CFL_{adv} = \frac{U \Delta t}{\Delta x} \implies \frac{1}{\Delta t^2 c^2} = \frac{M^2}{CFL_{adv}^2 \Delta x^2}$
- ▶ A finite centred difference discretization of the hyperbolic pressure equation

$$-\frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} + \nabla^2 p = 0$$

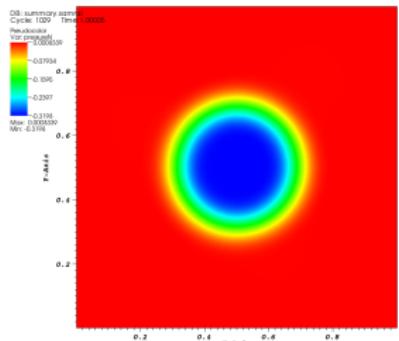
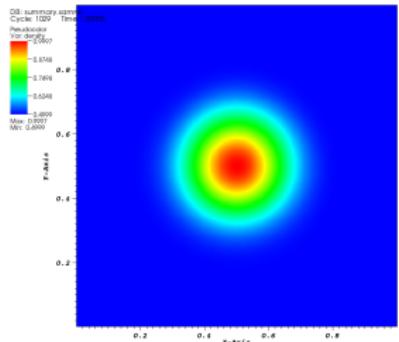
gives (in progress...)

$$-\frac{2}{c^2 \Delta t^2} (p^{n+1} - p^n) + \frac{1}{4} \nabla^2 p^{n+1} = \frac{2}{\Delta t} \rho^n \nabla \cdot \mathbf{v}^*$$

$T = 1\text{ s}$

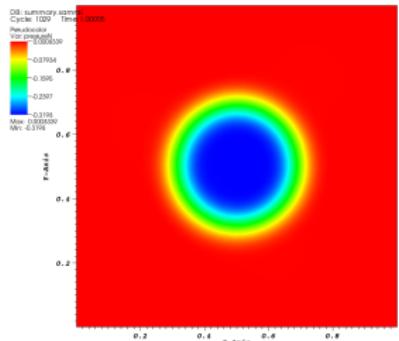
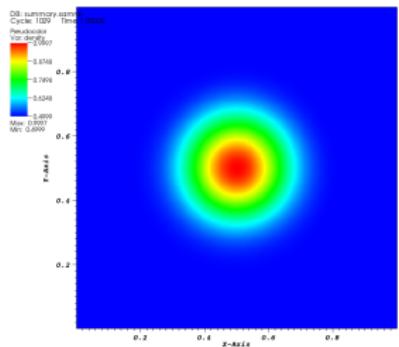


COMPRESSIBLE



$T = 1\text{ s}$

COMPRESSIBLE



PSEUDO-INCOMPRESSIBLE

